

* FERMI DIRAC STATISTICS:-

- > The Fermi Dirac distribution function gives the statistical behaviour of free electrons in metals and heavily doped semiconductors.
- > The particles that obey the Fermi Dirac statistics are sometimes called "Fermions".
- > For a system of N indistinguishable non-interacting particles obeying Pauli Exclusion principle with corresponding energies as E_1, E_2, \dots, E_N and the multiplicity or the degeneracy of the energy level as g_i .

The number of ways in which N_i particles are arranged in g_i quantum states is given by

$$g_i C_{N_i} = \frac{g_i!}{N_i! (g_i - N_i)!} \dots (1)$$

Each quantum states can accommodate only one particle in accordance with Pauli's Exclusion principle.

$$F = \prod_{i=1}^n \frac{g_i!}{N_i! (g_i - N_i)!} \dots (2)$$

F = no. of ways of distributing N_1, N_2, \dots, N_n particles energy levels

Π = product

Applying Stirling's theorem in the above equation:

$$\ln F = \sum_{i=1}^n \left[g_i (\ln g_i - 1) - N_i (\ln N_i - 1) - (g_i - N_i) \right]$$

$$\left\{ \ln(g_i - N_i) - (g_i - N_i) \right\}$$

$$= \sum_{i=1}^n \left[g_i \ln g_i - N_i \ln N_i - (g_i - N_i) \ln (g_i - N_i) \right] \dots (3)$$

Since the system is in equilibrium, F or $\ln F$ must be maximised. The total number of particles N and the total energy are constants.

i.e. $N = \sum_{i=1}^n N_i = \text{constant} \dots (4)$

$$U = \sum_{i=1}^n N_i E_i = \text{constant} \dots (5)$$

Applying Lagrange's method of undetermined multipliers to obtain (3), (4) and (5).

$$\ln(g_i - N_i) - \ln N_i + \alpha + \beta E_i = 0.$$

α, β are constants.

$$f(E_i) = \frac{N_i}{g_i} = \frac{1}{1 + e^{-\alpha - \beta E_i}} \dots (6)$$

This is the Fermi Dirac Distribution function.

$\beta = \frac{-1}{k_B T}$; where $k_B = \text{Boltzmann constant}$
 $T = \text{absolute temperature}$
 $\alpha = \frac{E_F}{k_B T}$, $E_F = \text{Fermi Energy}$.

The FD distribution function takes the form

$$f(E_i) = \frac{1}{1 + e^{(E_i - E_f)/k_B T}} \quad (7)$$

→ FD statistics is the probability that a quantum states of energy E_i is occupied.

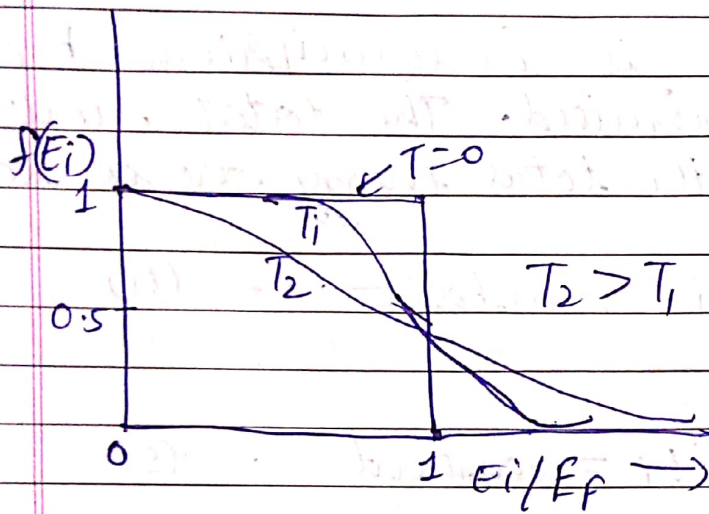


Fig: Plot of the Fermi distribution function $f(E_i)$ against normalized energy E_i/E_f at different temperatures

1) $T=0$, $f(E_i)=1$ for $E_i < E_f$ and $f(E_i)=0$ for $E_i > E_f$.

→ At absolute zero of temperature, the Fermi level represents the highest occupied energy levels

2) $T > 0$, $f(E_i)$ is close to unity for $E_i \ll E_f$ and approaches zero for $E_i \gg E_f$.

3) If the temperature is not very large $f(E_i)$ varies rapidly from about unity to about zero over an energy range of few times $k_B T$ around E_f .

- Thus ~~at~~ the Fermi level is that energy level for which the probability of occupation at $T > 0$ is $\frac{1}{2}$.
- At low temperatures when $f(E_i)$ is nearly a step function the distribution function is said to be strongly degenerate.
- At very high temperatures it is said to be nearly non-degenerate.

* BOSE - EINSTEIN STATISTICS.

- Bose-Einstein statistics gives the statistical behaviour of an ensemble of indistinguishable particles which do not follow Pauli Exclusion Principle. Examples are photons and phonons.
- The particles that obey the BE statistics are sometimes called Bosons.

The total number of ways V of distributing N_1, N_2, \dots, N_n particles in an energy level is the product of the terms represented by

$$M = \frac{(N_i + g_i - 1)!}{N_i! (g_i - 1)!}, \text{ where } g_i = \text{states} \dots (8)$$

is given by

$$V = \prod_{i=1}^n \frac{(N_i + g_i - 1)!}{N_i! (g_i - 1)!} \dots (9)$$

Applying Stirling's theorem in the above equation and then applying ~~log~~ Lagrange's method of undetermined multipliers as did in the FD distribution process we get

$$\sum_{i=1}^n [\ln(N_i + g_i - 1) - \ln N_i + \alpha + \beta E_i] S N_i = 0 \quad \dots (10)$$

$\alpha, \beta \rightarrow$ constants

The Bose Einstein distribution function yields

$$f(E_i) = \frac{N_i}{g_i} = \frac{1}{e^{-\alpha - \beta E_i} - 1} \quad \dots (11)$$

\rightarrow At very high temperatures, the particles are distributed over a wide energy range so that the number of particles in any energy range would be much smaller than the number of quantum states in that range i.e. $g_i \gg N_i$
 $\therefore \ln g_i - \ln N_i + \alpha + \beta E_i = 0 \quad \dots (12)$

The MB distribution function is then reduced to

$$f(E_i) = e^{\alpha + \beta E_i} \quad \dots (13)$$

$$\beta = \frac{-1}{k_B T} \quad \therefore f(E_i) = \frac{1}{e^{-\alpha} e^{E_i/k_B T} - 1}$$

At high temperatures α must be very large negative number. As the temperature approaches zero, $\alpha \rightarrow 0$.

$$f(E_i) = \frac{1}{e^{E_i/k_B T} - 1} \quad \dots (14)$$

\rightarrow At the absolute zero of temperature the particles tend to occupy the lowest energy state. This phenomenon is known as Bose-condensation.

* Comparison of Bose-Einstein and Fermi-Dirac distribution functions.

FERMI DIRAC

BOSE-EINSTEIN.

1) It holds for indistinguishable particles obeying the Pauli's Exclusion Principle

It holds for distinguishable particles not obeying Pauli's Exclusion Principle.

2) $f(E) = \frac{1}{e^{(E-E_F)/k_B T} + 1}$

$f(E) = \frac{1}{e^{-\alpha} e^{E/k_B T} - 1}$

3) It applies to electron gas in metals, protons, muons, neutrons, electrons, positrons etc

It applies to photon gas, phonon gas, pions, and the particles with integral or zero spin.