

Unit-2

Newton Forward And Backward Interpolation

Interpolation is the technique of estimating the value of a function for any intermediate value of the independent variable, while the process of computing the value of the function outside the given range is called **extrapolation**.

Forward Differences : The differences $y_1 - y_0, y_2 - y_1, y_3 - y_2, \dots, y_n - y_{n-1}$ when denoted by $\Delta y_0, \Delta y_1, \Delta y_2, \dots, \Delta y_{n-1}$ are respectively, called the first forward differences. Thus the first forward differences are :

Forward difference table

x	y	Δy	$\Delta^2 y$	$\Delta^3 y$	$\Delta^4 y$	$\Delta^5 y$
x_0	y_0					
x_1 $(= x_0 + h)$	y_1	Δy_0	$\Delta^2 y_0$			
x_2 $(= x_0 + 2h)$	y_2	Δy_1	$\Delta^2 y_1$	$\Delta^3 y_0$		
x_3 $= (x_0 + 3h)$	y_3	Δy_2	$\Delta^2 y_2$	$\Delta^3 y_1$	$\Delta^4 y_0$	
x_4 $= (x_0 + 4h)$	y_4	Δy_3	$\Delta^2 y_3$	$\Delta^3 y_2$	$\Delta^4 y_1$	$\Delta^5 y_0$
x_5 $= (x_0 + 5h)$	y_5	Δy_4				

NEWTON'S GREGORY FORWARD INTERPOLATION FORMULA :

This formula is particularly useful for interpolating the values of $f(x)$ near the beginning of the set of values given. h is called the interval of difference and $u = (x - a) / h$, Here a is first term.

θ°	45°	50°	55°	60°
$\sin \theta$	0.7071	0.7660	0.8192	0.8660

x°	Differences			
	10^1y	$10^1\Delta y$	$10^1\Delta^2y$	$10^1\Delta^3y$
45°	7071	589	-57	-7
50°	7660	532	-64	
55°	8192	468		
60°	8660			

Backward Differences : The differences $y_1 - y_0, y_2 - y_1, \dots, y_n - y_{n-1}$ when denoted by $\nabla y_1, \nabla y_2, \dots, \nabla y_n$, respectively, are called first backward difference. Thus the first backward differences are :

Backward difference table

x	y	∇y	$\nabla^2 y$	$\nabla^3 y$	$\nabla^4 y$	$\nabla^5 y$
x_0	y_0					
x_1 ($= x_0 + h$)	y_1	∇y_1				
x_2 ($= x_0 + 2h$)	y_2	∇y_2	$\nabla^2 y_2$	$\nabla^3 y_3$		
x_3 ($= x_0 + 3h$)	y_3	∇y_3	$\nabla^2 y_3$	$\nabla^3 y_4$	$\nabla^4 y_4$	
x_4 ($= x_0 + 4h$)	y_4	∇y_4	$\nabla^2 y_4$	$\nabla^3 y_5$	$\nabla^4 y_5$	$\nabla^5 y_5$
x_5 ($= x_0 + 5h$)	y_5	∇y_5	$\nabla^2 y_5$			

NEWTON'S GREGORY BACKWARD INTERPOLATION FORMULA :

This formula is useful when the value of $f(x)$ is required near the end of the table. h is called the interval of difference and $u = (x - a_n) / h$, Here a_n is last term.

Example :

Input : Population in 1925

<i>Year (x):</i>	<i>1891</i>	<i>1901</i>	<i>1911</i>	<i>1921</i>	<i>1931</i>
<i>Population (y):</i> <i>(in thousands)</i>	<i>46</i>	<i>66</i>	<i>81</i>	<i>93</i>	<i>101</i>

Output :

x	y	Vy	V^2y	V^3y	V^4y
1891	46	20			
1901	66	15	- 5		
1911	81	12	- 3	2	
1921	93	8	- 4	- 1	- 3
1931	101				

Value in 1925 is 96.8368

Stirling Interploation

Stirling Approximation or Stirling Interpolation Formula is an interpolation technique, which is used to obtain the value of a function at an intermediate point within the range of a discrete set of known data points .

Stirling Formula is obtained by taking the average or mean of the Gauss Forward and Gauss Backward Formula . Both the Gauss Forward and Backward formula are formulas for obtaining the value of the function near the middle of the tabulated set .

How to find

Stirling Approximation involves the use of forward difference table, which can be prepared from the given set of x and $f(x)$ or y as given below –

x	y	Δy	$\Delta^2 y$	$\Delta^3 y$	$\Delta^4 y$	$\Delta^5 y$
x_0	y_0					
x_1	y_1	Δy_0	$\Delta^2 y_0$			
x_2	y_2	Δy_1	$\Delta^2 y_1$	$\Delta^3 y_0$		
x_3	y_3	Δy_2	$\Delta^2 y_2$	$\Delta^3 y_1$	$\Delta^4 y_0$	$\Delta^5 y_0$
x_4	y_4	Δy_3	$\Delta^2 y_3$	$\Delta^3 y_2$	$\Delta^4 y_1$	
x_5	y_5	Δy_4				

This table is prepared with the help of x and its corresponding $f(x)$ or y . Then, each of the next column values is computed by calculating the difference between its preceeding and succeeding values in the previous column, like

$$\Delta y_0 = y_1 - y_0, \Delta y_1 = y_2 - y_1,$$

$$\Delta^2 y_0 = \Delta y_1 - \Delta y_0, \text{ and so on.}$$

Now, the Gauss Forward Formula for obtaining $f(x)$ or y at a is:

where, $p = (a - x_0)/h$,

a is the point where we have to determine $f(x)$, x is the selected value from the given

Δy

x which is closer to a (generally, a value from the middle of the table is selected), and h is the difference between any two consecutive x . Now, y becomes the value corresponding to x and values before x have negative subscript and those after have positive subscript, as shown in the table below –

x	y	Δy	$\Delta^2 y$	$\Delta^3 y$	$\Delta^4 y$
\vdots	\vdots				
x_{-2}	y_{-2}	Δy_{-2}			
x_{-1}	y_{-1}	Δy_{-1}	$\Delta^2 y_{-2}$		
x_0	y_0		$\Delta^2 y_{-1}$	$\Delta^3 y_{-2}$	
x_1	y_1	Δy_0	$\Delta^2 y_0$	$\Delta^3 y_{-1}$	$\Delta^4 y_{-2}$
x_2	y_2	Δy_1			
\vdots	\vdots				

Stirling's Formula gives a good approximation for $n!$ in terms of elementary functions. Before stating the formula, we introduce the following notation:

if $f(n)$ is a function and $g(n)$ is a function, then we write $f(n) \sim g(n) \leftrightarrow \lim_{n \rightarrow \infty} \frac{f(n)}{g(n)} = 1$. The statement $f(n) \sim g(n)$ is read $f(n)$ is asymptotic to $g(n)$ as $n \rightarrow \infty$.

For example, one verifies that $n^2 \sim (n+1)^2$ and $\sqrt{1+n} \sim \sqrt{n}$.

Here is Stirling's Formula: $n! \sim n^n e^{-n} \sqrt{2\pi n}$. The following graph shows a plot of the function $h(n) = \frac{n!}{n^n e^{-n} \sqrt{2\pi n}}$, confirming Stirling's Formula: $h(n) \rightarrow 1$ as $n \rightarrow \infty$. It turns out that $h(n)$ is decreasing so $n^n e^{-n} \sqrt{2\pi n}$ always underestimates $n!$ by a small amount.

10 20 30 40 50 60 70 1.03 1.04 1.02 1.01 Figure 1 :

Stirling's Formula The proof of Stirling's Formula is beyond the scope of this course. Instead of proving the formula, we rather give a proof of a weaker statement: we show that for every positive integer n , $n^n e^{-n} < n! < (n+1)^n e^{-n}$. (1) This does not prove Stirling's Formula, but it gives motivation for the $n^n e^{-n}$ term in the formula. The proof of the $\sqrt{2\pi n}$ part of the formula is more difficult.

1 First Proof — To prove (1), we just have to show (by taking logarithms): $n \log n - n < \log(n!) < (n+1) \log(n+1) - n$. Since $n! = n \cdot (n-1) \cdot (n-2) \cdots 2 \cdot 1$, $\log(n!) = \log 1 + \log 2 + \cdots + \log n$. The sum on the right can be estimated by integrals: let's show that $\log 1 + \log 2 + \cdots + \log n < \int_1^{n+1} \log x \, dx$. To see this, note that the integral represents the area under the curve $y = \log x$ (the red curve in the left plot below) for $1 \leq x \leq n+1$, whereas the sum $\log 1 + \log 2 + \cdots + \log n$ represents adding up the areas of rectangles with height $\log k$ for $k = 1, 2, \dots, n$ (see green step function in the left plot below). Now we can work out the integral: $\int_1^{n+1} \log x \, dx = x \log x - x + 1 \Big|_1^{n+1} = (n+1) \log(n+1) - n$. Therefore $\log(n!) < (n+1) \log(n+1) - n$. We're going to do the same thing to prove $\log(n!) > n \log n - n$: we claim that $\log 1 + \log 2 + \cdots + \log n > \int_0^n \log x \, dx$.

This is shown in the figure on the right, with the red curve representing $\log x$ and the rectangles representing $\log(1) + \log(2) + \cdots + \log(n)$.

Figure 2 : Approximating $\log(n!)$ Therefore $\log(n!) > n \log n - n$, which completes the proof of (1).

Bessel's Interpolation

Interpolation is the technique of estimating the value of a function for any intermediate value of the independent variable, while the process of computing the value of the function outside the given range is called **extrapolation**.

Central differences : The central difference operator δ is defined by the relations :

$$y_1 - y_0 = \delta y_{1/2}, y_2 - y_1 = \delta y_{3/2}, \dots, y_n - y_{n-1} = \delta y_{n - \frac{1}{2}}.$$

Similarly, high order central differences are defined as :

$$\delta y_{3/2} - \delta y_{1/2} = \delta^2 y_1, \quad \delta y_{5/2} - \delta y_{3/2} = \delta^2 y_2$$

Note – The central differences on the same horizontal line have the same suffix

Central difference table

x	y	δy	$\delta^2 y$	$\delta^3 y$	$\delta^4 y$	$\delta^5 y$
x_0	y_0	$\delta y_{1/2}$				
x_1	y_1	$\delta y_{3/2}$	$\delta^2 y_1$	$\delta^3 y_{3/2}$		
x_2	y_2	$\delta y_{5/2}$	$\delta^2 y_2$	$\delta^3 y_{5/2}$	$\delta^4 y_2$	
x_3	y_3	$\delta y_{7/2}$	$\delta^2 y_3$	$\delta^3 y_{7/2}$	$\delta^4 y_3$	$\delta^5 y_{5/2}$
x_4	y_4	$\delta y_{9/2}$	$\delta^2 y_4$			
x_5	y_5					

Bessel's Interpolation formula –

$$\begin{aligned}
 f(u) = & \left\{ \frac{f(0) + f(1)}{2} \right\} + \left(u - \frac{1}{2} \right) \Delta f(0) \\
 & + \frac{u(u-1)}{2!} \left\{ \frac{\Delta^2 f(-1) + \Delta^2 f(0)}{2} \right\} \\
 & + \frac{(u-1) \left(u - \frac{1}{2} \right) u}{3!} \Delta^3 f(-1) \\
 & + \frac{(u+1) u (u-1) (u-2)}{4!} \left\{ \frac{\Delta^4 f(-2) + \Delta^4 f(-1)}{2} \right\} + \dots
 \end{aligned}$$

It is very useful when $u = 1/2$. It gives a better estimate when $1/4 < u < 3/4$. Here $f(0)$ is the origin point usually taken to be mid point, since Bessel's is used to interpolate near the centre. h is called the interval of difference and $u = (x - f(0)) / h$. Here $f(0)$ is term at the origin chosen.

Examples –

Input	:		Value	at	27.4	?
x :	25	26	27	28	29	30
$f(x)$:	4.000	3.846	3.704	3.571	3.448	3.333

Output :

u	$10^3 f(u)$	$10^3 \Delta f(u)$	$10^3 \Delta^2 f(u)$	$10^3 \Delta^3 f(u)$	$10^3 \Delta^4 f(u)$	$10^3 \Delta^5 f(u)$
-2	4000	-154				
-1	3847	-142	12			
→ 0	3704	-133	9	-3	4	
1	3571	-123	10	1	-3	-7
→ 2	3448	-115	8	-2		
3	3333					

Value at 27.4 is 3.64968